

Ground-based characterization of (98943) 2001 CC₂₁, the target of Hayabusa2# space mission

Marcel M. Popescu¹, Eri Tatsumi², Javier Licandro³, Miguel R. Alarcon³, Javier Rodriguez Rodriguez⁴, Miquel Serra-Ricart⁵, Julia de Leon⁷, Joaquin Fernandez Martin⁵, David Morate⁶, Gabriel N. Simion⁷, Bogdan Alexandru Dumitru⁸, Daniel Nicolae Bertesteanu¹¹, George Pantelimon Prodan¹, Masatoshi Hirabayashi¹⁰

¹University of Craiova, Str. A. I. Cuza nr. 13, 200585 Craiova, Romania; ²Institute of Space and Astronautical Science (ISAS), Japan Aerospace Exploration Agency (JAXA), Sagamihara, Kanagawa, Japan; ³Instituto de Astrofísica de Canarias (IAC), C/Via Lactea s/n, 38205 La Laguna, Tenerife, Spain; ⁴Instituto Universitario de Ciencias y Tecnologías Espaciales de Asturias (ICTEA), University of Oviedo, C. Independencia 13, E-33004 Oviedo, Spain; ⁵Light Bridges S. L., Observatorio del Teide, Carretera del Observatorio S/N, E-38500 Guimar, Tenerife, Canarias, Spain; ⁶Centro de Estudios de Física del Cosmos de Aragón (CEFCA), Unidad Asociada al CSIC, Plaza San Juan 1, 44001, Spain; ⁷Astronomical Institute of the Romanian Academy, 5 Cușitul de Argint, 040557 Bucharest, Romania; ⁸Institute of Space Science (ISS) 409, Atomistilor Street, 077125 Magurele, Ilfov, Romania; ⁹Astroclubul Bucuresti Blvd Lascar Catargiu 21, 10663 Bucharest, Romania; ¹⁰Georgia Institute of Technology, Atlanta, GA 30320, United States

The near-Earth asteroid (98943) 2001 CC₂₁ is the target of the Hayabusa2 extended mission (Hayabusa2#). Ground-based telescope observations play an important role in providing key scientific data for this mission. A detailed study of the asteroid was conducted between 2022 and 2024. During this time frame it reached an apparent magnitude as bright as 16.5. We determined its rotation period to be $P = 5.021516 \pm 0.000106$ hours and its absolute magnitude $H = 18.693 \pm 0.10$. Based on these values, we estimated its diameter to be $D = 0.523 \pm 0.20$ km. The asteroid was classified as an Sq-type in the Bus-DeMeo taxonomy using a high signal-to-noise ratio spectrum, covering both the visible and near-infrared regions. Its mineral composition is likely similar to LL/L ordinary chondrites, with an olivine-to-pyroxene ratio ($ol/(ol+px)$) of 0.60, a fayalite (Fa) content of 28.5 mol%, and a ferrosilite (Fs) content of 23.4 mol%. Simultaneous observations in the g, r, i, and zs broadband filters revealed no significant large-scale heterogeneity on the surface of 2001 CC₂₁. The extensive lightcurve data allowed us to estimate the asteroid's convex shape and pole orientation as $\lambda = 301^\circ \pm 35^\circ$, $\beta = 89^\circ (+1^\circ/-6^\circ)$, and an axial tilt (obliquity) of $\epsilon = 5^\circ \pm 3^\circ$.

References

[1] Popescu, Tatsumi et al. 2024, PSJ under review