

## PROBING METEORITE MATRIX USING NON-DESTRUCTIVE TECHNIQUES.

E. Vaccaro<sup>1,2</sup>, S. S. Russell<sup>2</sup>, T. Goral<sup>1</sup>, J. Najorka. <sup>1</sup>Dept of Earth Sciences, The Natural History Museum, London, UK. E-mail: e.vaccaro@nhm.ac.uk. <sup>2</sup>Planetary & Space Sciences, Open University, Milton Keynes, UK.

**Introduction:** The minute quantities or particle size of samples returned from future space missions makes it vital to develop non-invasive diagnostic research methods to study these precious materials. The mineralogy of such samples is a key parameter to determine, but presents many challenges when material is limited or high spatial resolution is required. In this study we have carried out a combination of several non-destructive analysis of selected matrix regions of primitive carbonaceous chondrites ALHA77307 and QUE 99177 to investigate their nature.

**Experimental methods:** Ultra high-resolution image maps were acquired using a Carl Zeiss Ultra Plus Field Emission SEM. Image manipulation software “Image J” was used to estimate the clast and pore abundances by thresholding different grey-scale values from the image maps. Element maps of the same matrix regions were acquired, at 15kV and 5kV using a FEI Quanta 650 SEM with an innovative XFlash QUAD 5060F Bruker detector. [1]. Micro-XRD analyses were directly collected from the selected areas on the thin section using the Rigaku D max Rapid II. A pin-hole of 30  $\mu\text{m}$  was used to achieve an X-ray beam footprint on the sample of  $\sim 50 \times 500 \mu\text{m}$  [2].

**Results:** The micro-XRD patterns indicate the presence of olivine, clinoenstatite, pyrrhotite, magnetite and kamacite in the matrix of the meteorites studied. Clast and pore abundances in 5 image maps of each meteorite (each map having an area of  $\sim 20 \times 10 \mu\text{m}$ ) show the average abundance of metal grains in ALHA77307 and QUE 99177 is  $3 \pm 2.3\%$  and  $2 \pm 0.8\%$  respectively, the average porosity is  $6 \pm 4.1\%$  and  $3 \pm 1.6\%$  respectively, and the average abundance of crystalline silicates and oxides are  $65 \pm 12.1\%$  and  $46 \pm 4.8\%$  respectively. A groundmass of amorphous material accounts for  $26 \pm 16.2\%$  in ALHA77307 and  $49 \pm 4.0\%$  in QUE 99177. In both samples the element maps acquired show that the main cations of the amorphous groundmass are Al, Si and Fe, consistent with previous works e.g. [3], [4]. The matrix of ALHA77307 contains some acicular grains of Mg-silicate grains, a morphology indicative of condensation from solar nebular.

**Conclusions:** The phase identification using in situ micro-XRD demonstrates that this technique has the potential to be a useful tool in the characterization of minute samples or local areas of interest. The element maps acquired, using a FEI Quanta 650 SEM with an innovative Bruker detector, demonstrated to be capable of providing chemical information at spatial resolution down to a volume of 200 nm in depth by 100 nm in radius at 5kV. This has proved to be a useful tool for non-destructive chemical analysis at high spatial resolution.

**References:** [1] E. Vaccaro et al., Abstract #5327. *MetSoc 2014* [2] E. Vaccaro et al., Abstract #5348. *MetSoc 2014*. [3] A. Brearley, 1993, *Geochim. et Cosmochim. Acta* 57. [4] M. Neyda, A. Brearley, 2010, *Geochim. et Cosmo-chim. Acta* 74.