

### H, C and N isotopic compositions of NaCl bearing organic sample in Hayabusa category 3.

M. Ito<sup>1</sup>, M. Uesugi<sup>2</sup>, H. Naraoka<sup>3</sup>, H. Yabuta<sup>4</sup>, F. Kitajima<sup>3</sup>, H. Mita<sup>5</sup>, Y. Takano<sup>1</sup>, Y. Karouji<sup>2</sup>, T. Yada<sup>2</sup>, Y. Ishibashi<sup>2</sup>, T. Okada<sup>2</sup> and M. Abe<sup>2</sup>. <sup>1</sup>JAMSTEC. [motoo@jamstec.go.jp](mailto:motoo@jamstec.go.jp). <sup>2</sup>JAXA/ISAS. <sup>3</sup>Kyushu University. <sup>4</sup>Osaka University. <sup>5</sup>Fukuoka Inst. Tech.

We have reported H, C and N isotopic compositions of Hayabusa category 3 samples, RB-QD04-0047-02, RA-QD02-0120, and RB-QD04-0001 [1]. All samples show terrestrial H, C, and N isotopic compositions within errors, and none of these samples contain  $\mu\text{m}$ -sized hot spots with anomalous H, C, and N isotopic compositions, unlike previous isotope studies for extraterrestrial organic materials of IOM [2], nano-globules in chondrites [3], IDPs [4], and STARDUST cometary dusts [5, 6]. It is difficult to conclude, based on the isotope data, whether these Hayabusa category 3 samples are terrestrial contaminants or extraterrestrial materials. We have, thus, continued investigations of category 3 samples through a developed sequential analyses including ToF-SIMS, NanoSIMS, (S)TEM, FT-IR, Raman spectroscopy and XANES that described in [7].

In this study, new category 3 samples of RB-QD04-0037-01 and RA-QD02-0180 were available for isotopic measurements with a NanoSIMS ion microprobe. Both samples were mainly composed of C, N and O based on FE-SEM with EDS analysis at JAXA/ESCuC [8]. It is noted that RA-QD02-0180 contains trace amount of Na, K and Cl [7].

These samples show homogeneous and terrestrial H, C and N isotopic composition ( $\delta\text{D} = -8 \sim +2 \text{‰}$ ,  $\delta^{13}\text{C} = -8 \sim 0 \text{‰}$  and  $\delta^{15}\text{N} = +5 \sim +25 \text{‰}$  for RB-QD04-0037-01;  $\delta\text{D} = +1 \sim +13 \text{‰}$ ,  $\delta^{13}\text{C} = -21 \sim -29 \text{‰}$  and  $\delta^{15}\text{N} = +13 \sim +14 \text{‰}$  for RA-QD02-0180). In RA-QD02-0180, we found numerous O, S and Cl enriched  $\mu\text{m}$ -sized regions that scattered among the sample. Locations of O, S and Cl areas were different so these element distributions were not related each other. We readily determined C and N isotopic ratios of each of the  $^{16}\text{O}$ ,  $^{32}\text{S}$  and  $^{35}\text{Cl}$  areas in the field of view by defining regions of interest. Based on isotope images of  $\delta\text{D}$ ,  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$ , there were no obvious isotopic anomalies from any of  $^{16}\text{O}$ ,  $^{32}\text{S}$  and  $^{35}\text{Cl}$  areas in the sample. Bridge et al. [9] reported that halite assemblage was found in the Nakhla Martian meteorite implying a low-temperature formation process by crystallisation from an aqueous fluid. It is important to explore the nature and origin of the NaCl bearing sample for understanding of water-mineral interaction and relationship with organics. We, therefore, continue to investigate RA-QD02-0180 sample by coordinated analysis with NanoSIMS, TEM, FIB, C-XANES, Raman spectrometry and FT-IR.

**References:** [1] Ito et al. 2014. *Earth Planets Space* 66:91. [2] Busemann et al. 2006. *Science* 312:727. [3] Nakamura-Messenger et al. 2006. *Science* 314:1439. [4] Messenger 2000. *Nature* 404:968. [5] McKeegan et al. 2006. *Science* 314:1724. [6] Sanford et al. 2010. *Science* 314:1720. [7] Uesugi et al. 2014. *Earth Planets Space* 66:102. [8] Hayabusa sample catalogue: <http://hayabusaa.isas.jaxa.jp/catalog/cat3>. [9] Bridges et al. 1999. *MaPS* 34:407.